



**HAL**  
open science

# First-principles calculations of pressure broadening for the case of N 2 -perturbed 118 GHz fine-structure line in O 2 (X 3 Σ -g )

Maciej Gancewski, Hubert Józwiak, E. Quintas-Sánchez, R. Dawes, Franck Thibault, P Wcislo

► **To cite this version:**

Maciej Gancewski, Hubert Józwiak, E. Quintas-Sánchez, R. Dawes, Franck Thibault, et al.. First-principles calculations of pressure broadening for the case of N 2 -perturbed 118 GHz fine-structure line in O 2 (X 3 Σ -g ). 25th International Conference on Spectral Line Shapes, ICSSL 2022, Jun 2022, Caserta, Italy. hal-03718884

**HAL Id: hal-03718884**

<https://univ-rennes.hal.science/hal-03718884>

Submitted on 19 Sep 2022

**HAL** is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers.

L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

# First-principles calculations of pressure broadening for the case of N<sub>2</sub>-perturbed 118 GHz fine-structure line in O<sub>2</sub>(X<sup>3</sup>Σ<sub>g</sub><sup>-</sup>)

M. Gancewski<sup>1</sup>, H. Jóźwiak<sup>1</sup>, E. Quintas-Sánchez<sup>2</sup>, R. Dawes<sup>2</sup>, F. Thibault<sup>3</sup>, P. Wcisło<sup>1</sup>

1. Institute of Physics, Faculty of Physics, Astronomy and Informatics, Nicolaus Copernicus University in Toruń, ul. Grudziądzka 5, 87-100 Toruń, Poland

2. Department of Chemistry, Missouri University of Science and Technology, Rolla, Missouri 65409-0010, USA

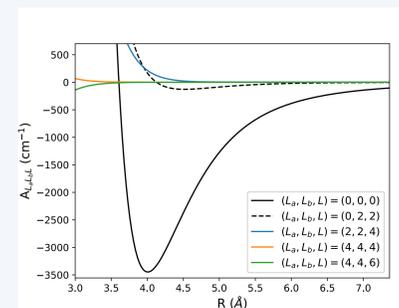
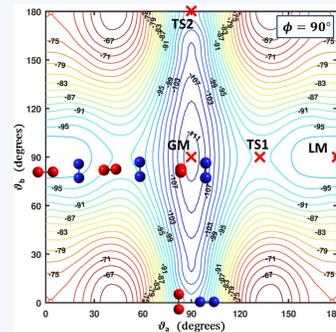
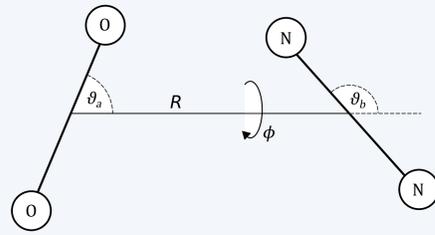
3. Univ. Rennes, CNRS, IPR (Institut de Physique de Rennes)-UMR 6251, Rennes F-35000, France

## What is this about?

- The general problem: *How is the absorption (emission) spectrum of O<sub>2</sub> modified when it is immersed in the large bath of N<sub>2</sub> molecules?*
- This model question is important in general considerations of atmospheric physics (~78% of N<sub>2</sub> & ~21% of O<sub>2</sub> in our atmosphere...)
- We want to attack this problem *ab initio* style, using methods of theoretical physics and quantum chemistry
- Many spectral lines & bands of O<sub>2</sub> will be affected by its interactions with N<sub>2</sub> in a complicated manner
- A thorough study of these is very challenging (preparation of the potential, solving the Schrödinger & kinetic equations...)
- We thus begin this program by addressing a simpler problem: *How much a single isolated line of O<sub>2</sub> is broadened by collisions with N<sub>2</sub> at different gas temperatures?*
- Here, we consider both (rigid) molecules to be in their ground states & we focus on the N<sub>2</sub>-induced perturbations of the isolated fine-structure line (centered at ~118 GHz) of O<sub>2</sub>(<sup>3</sup>Σ)

## Our methodology

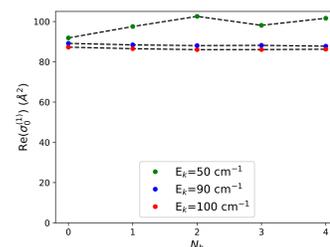
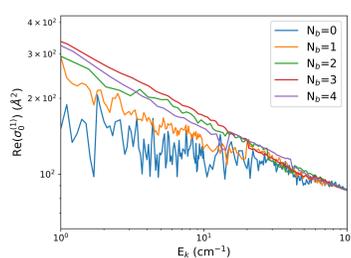
- In O<sub>2</sub>(X<sup>3</sup>Σ) each rotational level splits into three fine-structure levels due to spin-spin interaction; we study the broadening of the  $j=1 \rightarrow j=0$  fine-structure transition, occurring within the first rotational level
- By using angular momentum recoupling techniques [1, 2, 3] we separate the O<sub>2</sub> spin from the problem, and solve the Schrödinger equation for the collision of two spin-free particles on the 4-dimensional potential energy surface (PES)



- Our new 4-dimensional O<sub>2</sub>-N<sub>2</sub> PES is computed using the (AE)UCCSD(T)-F12b/CBS methodology with aid of the "AUTOSURF" code [1, 4]; the Schrödinger equation is solved numerically using our "BIGOS" quantum scattering code [5]
- We transform the obtained spin-free scattering matrices back into the representation which include the quantum spin number – we are thus in a position to compute the line-broadening for a purely fine-structure transition in the single rotational level
- The "BIGOS" code uses the transformed S-matrices to calculate the pressure broadening cross-section for the N<sub>2</sub> perturbation of the  $j=1 \rightarrow j=0$  fine-structure transition in O<sub>2</sub>
- The energy-averaged cross sections yield the pressure broadening parameter for this line (*i.e.*, the half-width at half maximum, HWHM) as a function of the gas temperature [1]

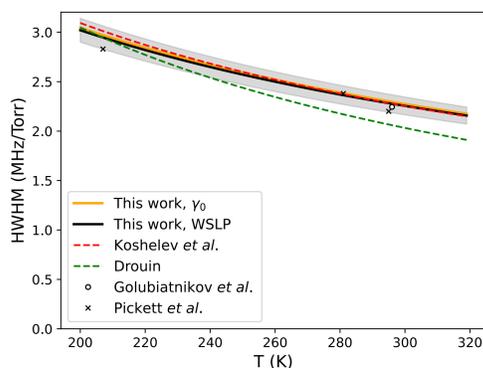
## The results

- We calculated the broadening cross sections in the kinetic energy range from 1 cm<sup>-1</sup> to ~100 cm<sup>-1</sup>
- Here, cross sections are given for first 5 rotational levels of N<sub>2</sub> (labeled with N<sub>b</sub>) as a function of kinetic energy; at high enough E<sub>k</sub> they become independent of the rotational state of the thermal bath (see the figure on the right)



- We extrapolate the cross sections given in the figure into the higher kinetic energies (up to ~4 000 cm<sup>-1</sup>); we then assume that for N<sub>2</sub> rotational states with N<sub>b</sub>>4, the cross sections are the same as for the N<sub>b</sub>=4; thus in the evaluation of the statistical partition sums over bath variables we are able to include the N<sub>2</sub> levels up to N<sub>b</sub>=31

- By suitable energy-averaging of the cross sections we obtain both the speed-independent & speed-dependent N<sub>2</sub>-broadening coefficients for the O<sub>2</sub> 118 GHz line. The temperature dependence of the speed-dependent HWHM is obtained by averaging the speed-dependent Lorentz profiles over the Maxwellian distribution



$$\gamma(T) = \gamma(T_0) \left( \frac{T_0}{T} \right)^{n_\gamma}$$

Experiment	T (K)	$\gamma$ (MHz/Torr)	$n_\gamma$
Pickett <i>et al.</i> <sup>19</sup>	207	2.83(6)	...
	281	2.38(5)	...
	295	2.2(1)	...
Golubiatnikov <i>et al.</i> <sup>20</sup>	297	2.245(20)	...
Drouin <sup>21</sup>	296	2.06(6)	1.00(5)
Makarov <i>et al.</i> <sup>15</sup>	300	2.255(14)	0.79(5)
Koshelev <i>et al.</i> <sup>16</sup>	296	2.278(2)	0.781(6)
This work, WSLP	296	2.28(9)	0.7231(1)
This work, $\gamma_0$	296	2.30(9)	0.7245(6)

- Our *ab initio* calculations agree with the experimental data (see [1] for the relevant papers) within the estimated 4% uncertainty of numerical calculations (shaded area). This paves the way for more elaborate future analyses of the general line-shape problem for atmospheric O<sub>2</sub> perturbations

## Bibliography

- Gancewski, M. *et al.* J. Chem. Phys. **155**, 124307 (2021)
- Offer, A. R. *et al.* J. Chem. Phys. **100**, 362 (1994)
- Corey, G. C. *et al.* J. Phys. Chem. **87**, 2723 (1983)
- Quintas-Sánchez, E. *et al.* J. Chem. Inf. Model. **59**, 262 (2018)
- Jóźwiak, H. *et al.* "BIGOS computer code", to be published

## Acknowledgements

M.G., H.J., and P.W. were supported by the National Science Centre in Poland through Project No. 2018/31/B/ST2/00720. R.D. and E.Q.-S. were supported by the U.S. Department of Energy (Award No. DE-SC0019740). This project was supported by the French-Polish PHC Polonium program (Project No. 42769ZK for the French part). This project was co-financed by the Polish National Agency for Academic Exchange under the PHC Polonium program (Grant No. dec. PPN/X/PS/318/2018). This research is a part of the program of the National Laboratory FAMO in Toruń, Poland.